

# PHYSICAL REVIEW LETTERS

VOLUME 83

30 AUGUST 1999

NUMBER 9

## The Transition Temperature of the Dilute Interacting Bose Gas

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(Received 1 June 1999)

We show that the critical temperature of a uniform dilute Bose gas increases linearly with the  $s$ -wave scattering length describing the repulsion between the particles. Because of infrared divergences, the magnitude of the shift cannot be obtained from perturbation theory, even in the weak coupling regime; rather, it is proportional to the size of the critical region in momentum space. By means of a self-consistent calculation of the quasiparticle spectrum at low momenta at the transition, we find an estimate of the effect in reasonable agreement with numerical simulations.

PACS numbers: 03.75.Fi

Determination of the effect of repulsive interactions on the transition temperature of a homogeneous dilute Bose gas at fixed density has had a long and controversial history [1–5]. While [1] predicted that the first change in the transition temperature,  $T_c$ , is of order the scattering length  $a$  for the interaction between the particles, neither the sign of the effect nor its dependence on  $a$  has been obvious. Recent renormalization group studies [4] predict an increase of the critical temperature. Numerical calculations by Grüter, Ceperley, and Laloë [6], and more recently by Holzmann and Krauth [7], of the effect of interactions on the Bose-Einstein condensation transition in a uniform gas of hard sphere bosons, and approximate analytic calculations by Holzmann, Grüter, and Laloë of the dilute limit [8], have shown that the transition temperature,  $T_c$ , initially rises linearly with  $a$ . The effect arises physically from the change in the energy of low momentum particles near  $T_c$  [8]. Here we analyze the leading order behavior of diagrammatic perturbation theory, and argue that  $T_c$  increases linearly with  $a$ . We then construct an approximate self-consistent solution of the single particle spectrum at  $T_c$  which demonstrates the change in the low momentum spectrum, and which enables us to calculate the change in  $T_c$  quantitatively.

We consider a uniform system of identical bosons of mass  $m$ , at temperature  $T$  close to  $T_c$  and use finite tem-

perature quantum many-body perturbation theory. We assume that the range of the two-body potential is small compared to the interparticle distance  $n^{-1/3}$ , so that the potential can be taken to act locally and be characterized entirely by the  $s$ -wave scattering length  $a$ . Thus we work in the limit  $a \ll \lambda$ , where  $\lambda = (2\pi\hbar^2/mk_B T)^{1/2}$  is the thermal wavelength. (We generally use units  $\hbar = k_B = 1$ .)

To compute the effects of the interactions on  $T_c$ , we write the density  $n$  as a sum over Matsubara frequencies  $\omega_\nu = 2\pi i\nu T$  ( $\nu = 0, \pm 1, \pm 2, \dots$ ) of the single particle Green's function,  $G(k, z)$ :

$$n = -T \sum_\nu \int \frac{d^3k}{(2\pi)^3} G(k, \omega_\nu), \quad (1)$$

where

$$G^{-1}(k, z) = z + \mu - \frac{k^2}{2m} - \Sigma(k, z), \quad (2)$$

with  $\mu$  the chemical potential. The Bose-Einstein condensation transition is determined by the point where  $G^{-1}(0, 0) = 0$ , i.e., where  $\Sigma(0, 0) = \mu$ .

The first effect of interactions on  $\Sigma$  is a mean field term  $\Sigma_{\text{mf}} = 2gn$ , where  $g = 4\pi\hbar^2 a/m$ ; the factor of 2 comes from including the exchange term. Such a contribution, independent of  $k$  and  $z$ , has no effect on the transition temperature, as it can be simply absorbed in a redefinition

of the chemical potential. To avoid carrying along such trivial contributions we define:

$$\frac{\hbar^2}{2m\zeta^2} = -(\mu - 2gn). \quad (3)$$

The quantity  $\zeta$  may be interpreted as the mean field correlation length. In the mean field approximation,  $\zeta$  becomes infinite at  $T_c$ ; however, in general, it remains finite, and functions here as an infrared cutoff.

Because the effects of interactions are weak, one could imagine calculating the change in  $T_c$  in perturbation theory. However, such calculations are plagued by infrared divergences. Power counting arguments reveal that the leading contribution to the self-energy,  $\Sigma(k \ll \zeta^{-1}, 0)$ , of a diagram of order  $a^n$  has the form:

$$\Sigma_n \sim T \left( \frac{a}{\lambda} \right)^2 \left( \frac{a\zeta}{\lambda^2} \right)^{n-2}. \quad (4)$$

In perturbation theory about the mean field, with the mean field criterion for the phase transition,  $\zeta \rightarrow \infty$  at  $T_c$ , all  $\Sigma_n$  diverge, starting with a logarithmic divergence at  $n = 2$ . More generally, the approach of  $\zeta$  towards  $\lambda^2/a$  in magnitude signals, according to the Ginsburg criterion, the onset of the critical region. Beyond, perturbation theory breaks down, since all  $\Sigma_n$  in Eq. (4) are of the same order of magnitude.

Even though the theory is infrared divergent, we can isolate the leading correction to the change in  $T_c$ , which, as we show, is of order  $a$ . Since the infrared divergences occur only in terms with zero Matsubara frequencies we separate, in Eq. (1), the contribution of the  $\nu = 0$  terms, writing

$$n(a, T) = -T \int \frac{d^3k}{(2\pi)^3} [G_{\nu=0}(k) + G_{\nu \neq 0}(k)], \quad (5)$$

where  $G_{\nu \neq 0}(k)$  is the sum of terms with  $\nu \neq 0$ . Similarly, the density of a noninteracting system with condensation temperature  $T$  is given by

$$\begin{aligned} n^0(T) &= -T \int \frac{d^3k}{(2\pi)^3} [G_{\nu=0}^0(k) + G_{\nu \neq 0}^0(k)] \\ &= \int \frac{d^3k}{(2\pi)^3} \frac{1}{e^{k^2/2mT} - 1} = \frac{\zeta(3/2)}{\lambda^3}, \end{aligned} \quad (6)$$

where  $\zeta(3/2) = 2.612$ . Since the nonzero Matsubara frequencies regularize the infrared behavior of the momentum integrals, the dependence of the  $\nu \neq 0$  terms in Eq. (5) on  $a$  is nonsingular at  $T_c$ . These terms, of order  $a^2$  at least, can be neglected. Thus to order  $a$ ,

$$n(a, T_c) - n^0(T_c) = -T_c \int \frac{d^3k}{(2\pi)^3} [G_0(k) - G_0^0(k)]. \quad (7)$$

To calculate the change in  $T_c$  at fixed density we equate  $n(a, T_c)$  at  $T_c$  with  $n^0(T_c^0)$  at the free particle transition temperature  $T_c^0$  and observe that  $n^0(T_c) = (T_c/T_c^0)^{3/2} n^0(T_c^0)$ ; thus in lowest order the change in

transition temperature  $\Delta T_c = T_c - T_c^0$  is given by

$$\frac{3}{2} \frac{\Delta T_c}{T_c} n^0(T_c^0) = T_c \int \frac{d^3k}{(2\pi)^3} [G_{\nu=0}(k) - G_{\nu=0}^0(k)], \quad (8)$$

where  $\Delta T_c = T_c - T_c^0$ . Thus

$$\frac{\Delta T_c}{T_c} = \frac{4\lambda}{3\pi\zeta(3/2)} \int_0^\infty dk \frac{U(k)}{k^2 + U(k)}, \quad (9)$$

where  $U(k) \equiv 2m[\Sigma_{\nu=0}(k) - \mu]$ .

Equation (9) for the leading correction to the critical temperature is crucial. The criterion for spatially uniform condensation is that  $U(0) = 0$ ; above the transition,  $U(0) > 0$ . At the transition,  $k^2 + U(k) > 0$  for  $k > 0$ . At large wave numbers,  $U \rightarrow 1/\zeta^2 > 0$ , and in the critical region, as we discuss below,  $U$  is also positive. Although we have not proved it rigorously, numerical simulations indicate that  $U$  is generally positive for  $k > 0$ , which implies that the integral in Eq. (9) and hence  $\Delta T_c$  is positive.

In the critical region,  $k < k_c$ , where  $k_c$  defines the scale of the critical region in momentum space,  $G_{\nu=0}$  has the scaling form [9]  $G_{\nu=0}^{-1}(k) = -k^{2-\eta} k_c^\eta F(k\xi)$ ;  $\xi$  is the coherence length which diverges at  $T_c$  as  $|T - T_c|^{-\nu}$ , and  $F$  is a dimensionless function, with  $F(\infty) \sim 1$ . The critical index,  $\eta$ , is given to leading order in the  $\epsilon = 4 - d$  expansion by  $\epsilon^2/54$  [10]. At  $T_c$ ,  $G_{\nu=0}^{-1}(k) \sim -k^{2-\eta} k_c^\eta$ , and thus  $U \sim +k^{2-\eta}$ . Both terms in Eq. (8) give a contribution of order  $k_c$ , so that  $\Delta T_c/T_c \sim k_c$ . As we shall see,  $k_c \sim a/\lambda^2$ , and hence  $\Delta T_c/T_c \sim a/\lambda$ .

To study the leading behavior in  $a$  quantitatively, we need concentrate only on the  $\nu = 0$  sector where the full finite temperature theory reduces to a classical field theory [10] defined by the action:

$$\begin{aligned} S\{\phi(r)\} &= \frac{1}{2mT} \int d^3r \left[ \nabla\phi^*(r) \cdot \nabla\phi(r) + \frac{1}{\zeta^2} |\phi(r)|^2 \right. \\ &\quad \left. + 4\pi a (|\phi(r)|^2 - \langle |\phi(r)|^2 \rangle)^2 \right]; \end{aligned} \quad (10)$$

the probability of a given field configuration entering the computation of expectation values is proportional to  $e^{-S\{\phi(r)\}}$ .

The classical theory is ultraviolet divergent, but superrenormalizable. The divergences appear only in the two-loop self-energy,  $\Sigma_{\nu=0}^{(2)}$ —effectively the second order self-energy written in terms of the full  $G_{\nu=0}$  rather than the zeroth order Green's functions—and can be removed by simple renormalization of the mean field coherence length,  $\zeta$ . Since, henceforth, we calculate only in the classical theory, we drop the subscript  $\nu = 0$ . The second order self-energy is

$$\Sigma(k) = -2g^2 \int \frac{d^3q}{(2\pi)^3} B(q) \frac{T}{\epsilon_{\mathbf{k}-\mathbf{q}}}, \quad (11)$$

where  $\epsilon_k = (k^2 + \zeta^{-2})/2m$ , and the ( $\nu = 0$ ) particle-hole bubble,

$$B(q) = \int \frac{d^3 p}{(2\pi)^3} \frac{T}{\epsilon_p \epsilon_{\mathbf{p}+\mathbf{q}}}, \quad (12)$$

is given by

$$B(q) = \frac{2\pi^2 \zeta}{T \lambda^4} b(\zeta q); \quad (13)$$

$b(x) \rightarrow 1/x$  for  $x \gg 1$  and  $b(0) = 1/\pi$ .

The integral in Eq. (11) is logarithmically divergent in the ultraviolet. But in the full theory the momentum integrals are cut off by distribution functions,  $f = (e^{k^2/2mT} - 1)^{-1}$ , and the ultraviolet behavior is regular. To control this divergence we introduce an ultraviolet momentum cut-off,  $\Lambda$ , in the classical theory, recognizing that it is in fact effectively determined in the full theory. Then

$$2m\Sigma(k) = -32\pi^2 \frac{a^2}{\lambda^4} \int_0^{\Lambda\zeta} dx x b(x) L(k\zeta, x), \quad (14)$$

where

$$L(k\zeta, x) = \frac{1}{k\zeta} \ln \left[ \frac{(x + k\zeta)^2 + 1}{(x - k\zeta)^2 + 1} \right]. \quad (15)$$

The divergent part of the integral comes from the large  $x$  tail of  $b(x)$ , and contributes  $-128(a/\lambda^2)^2 \ln(\Lambda\zeta)$  to  $2m\Sigma$ .

More generally we carry out a diagrammatic expansion of  $\Sigma$  in terms of the *self-consistent*  $\nu = 0$  Green's function, defined by  $2mG^{-1}(k) = -k^2 + \zeta^{-2} + 2m\Sigma(k, a, G, \Lambda)$ . Note that the dependence of  $\Sigma$  on  $\zeta$  enters only through the dependence of  $\Sigma$  on  $G$ . We define a *renormalized* mean field coherence length by

$$\frac{1}{\zeta_R^2} = \frac{1}{\zeta^2} - 128 \left( \frac{a}{\lambda^2} \right)^2 \ln(\Lambda\zeta_R). \quad (16)$$

Then  $G^{-1}(k)$  is given by

$$-2mG^{-1}(k) = k^2 + \zeta_R^{-2} + 2m\Sigma_F(k, a, G), \quad (17)$$

where

$$\Sigma_F(k, a, G) = \Sigma(k, a, G, \Lambda) + 128 \left( \frac{a}{\lambda^2} \right)^2 \ln(\Lambda\zeta_R) \quad (18)$$

is independent of  $\Lambda$ . As a function of  $\zeta_R$ , the Green's function is independent of the cutoff.

In fact, a simple power counting argument shows that the finite part of the self-energy has the form

$$\Sigma_F(k, a, G) = \frac{1}{2m\zeta_R^2} \sigma(k\zeta_R, J), \quad (19)$$

where

$$J = a\zeta_R/\lambda^2. \quad (20)$$

To see this structure we note that a term in the self-energy of order  $a^n$  is the product of a dimensionless function of  $k\zeta_R$  times the  $\Sigma_n$  of Eq. (4), with  $\zeta$  replaced by  $\zeta_R$  [11].

The criterion for condensation,  $\zeta_R^{-2} + 2m\Sigma_F(0, a, G) = 0$ , implies that

$$1 + \sigma(0, J) = 0. \quad (21)$$

Since  $\sigma(0)$  is a well-behaved function of only the parameter  $J$ , Eq. (21) determines the critical value of  $J = J^*$  for condensation, a dimensionless number independent of the parameters of the original problem. At condensation, the renormalized mean field coherence length  $\zeta_R$  tends to infinity as  $a \rightarrow 0$ , with the product  $a\zeta_R$  fixed, thus preventing a perturbative expansion in  $a$ .

At condensation  $U(k) = [\sigma(k\zeta_R, J^*) + 1]/\zeta_R^2$ , and Eq. (9) implies the change in  $T_c$

$$\frac{\Delta T_c}{T_c} = \frac{a}{\lambda} \left[ \frac{4}{3\pi\zeta(3/2)} \frac{1}{J^*} \int_0^\infty dx \frac{\sigma(x, J^*) + 1}{x^2 + \sigma(x, J^*) + 1} \right]. \quad (22)$$

Since  $J^*$  is determined by the condition (21), the result for  $\Delta T_c/T_c$  is linear and expected to be positive in  $a/\lambda$ .

We turn now to calculating  $\Delta T_c$  explicitly within a simple self-consistent model based on taking only the zero frequency component of the leading two-loop approximation self-energy, given by Eq. (11). We construct the  $\epsilon_p$  as self-consistent quasiparticle energies at the transition, i.e., solutions of the equation:

$$G^{-1}(k, \epsilon_k) = 0 = \epsilon_k - \frac{k^2}{2m} - [\Sigma(k) - \Sigma(0)]. \quad (23)$$

The low momentum behavior of  $\epsilon_k$  is determined by a familiar argument [12]. In order that the integral (11) converge in the infrared limit,  $\epsilon_k$  must behave, modulo possible logarithmic corrections, as  $\sim k^\alpha$ , where  $\alpha < 2$ . In this case, the term  $k^2/2m$  in (23) can be neglected at small  $k$ . We then expand  $\Sigma(k)$  about  $k = 0$ . For  $1 \leq \alpha < 4/3$  the self-energy is sufficiently convergent that  $\Sigma(k) - \Sigma(0) \sim k^2$  at small  $k$ , and thus cannot be the correct self-consistent solution. For  $\alpha$  with  $4/3 < \alpha < 2$  one has  $\Sigma(k) - \Sigma(0) \sim +k^{6-3\alpha}$ , so we find self-consistency,  $\Sigma(k) - \Sigma(0) \sim k^\alpha$ , for  $\alpha = 3/2$ . We write the small  $k$  part of the spectrum as

$$\epsilon_k = k_c^{1/2} k^{3/2} / 2m. \quad (24)$$

Here  $k_c$  is the wave vector around which the  $k^{3/2}$  at low  $k$  crosses over to the  $k^2/2m$  free-particle behavior.

To extract the low momentum structure, below the scale  $k_c$ , we evaluate the most divergent terms of

$$\Sigma(k) - \Sigma(0) = -2g^2 T \int \frac{d^3 q}{(2\pi)^3} B(q) \left( \frac{1}{\epsilon_{\vec{k}-\vec{q}}} - \frac{1}{\epsilon_q} \right); \quad (25)$$

at small  $k$ . Since the  $k^{3/2}$  structure arises from the small  $q$  behavior of the integral; we evaluate the bubble  $B(q)$ , Eq. (12), at small  $q$  with the spectrum (24) for  $k < k_c$  and  $k^2/2m$  for  $k > k_c$ . Then

$$B(q) = \frac{4m}{\pi\lambda^2 k_c} [\ln(k_c/q) + c], \quad (26)$$

with  $c \approx 2 + 2 \ln 2 - \pi/2 = 1.816$ . Thus,

$$\Sigma(k) - \Sigma(0) = \frac{1024\pi}{15m} \left(\frac{a}{\lambda^2}\right)^2 \left(\frac{k}{k_c}\right)^{3/2}. \quad (27)$$

Identifying the right side of Eq. (27) with  $k_c^{1/2}k^{3/2}/2m$ , we derive

$$k_c = 32 \left(\frac{2\pi}{15}\right)^{1/2} \frac{a}{\lambda^2} \approx 20.7 \frac{a}{\lambda^2}. \quad (28)$$

As expected, the scale of the unusual low momentum structure is  $a/\lambda^2$ .

Let us, for a first quantitative estimate, assume a spectrum at  $T_c$  of the form  $\epsilon_k = k_c^{1/2}k^{3/2}/2m$  for  $k \ll k_c$ , and  $(k^2 + k_c^2)/2m$  for  $k \gg k_c$ . We smoothly interpolate between these limits, writing  $U(k) = k_c^{1/2}k^{3/2}/[1 + (k/k_c)^{3/2}]$ . Thus  $\int dk U/(k^2 + U) \approx 1.2k_c$ , so that with Eq. (28),

$$\frac{\Delta T_c}{T_c} \approx 2.9an^{1/3}. \quad (29)$$

By comparison, Grüter, Ceperley, and Laloë [6] find  $\Delta T_c/T_c \approx 0.34an^{1/3}$ , while the more recent calculation of Holzmann and Krauth yields  $\Delta T_c/T_c \approx (2.2 \pm 0.2)an^{1/3}$ . The agreement of the numerical coefficient, given the simplicity of the approximations in evaluating the effect of interactions on the transition temperature, is satisfying. As will be reported in a fuller paper [14], this estimate agrees with that derived from the numerical self-consistent solution of Eq. (23).

The lowest two-loop calculation does not account fully for the modification of the transition temperature; indeed, at the critical point, all diagrams become comparable [13,14]. Consider, for example, summing the bubbles describing the repeated scattering of the particle-hole pair in  $B$  [15], thus replacing  $B$  in Eq. (11) by

$$B_{\text{eff}}(q) = \frac{B(q)}{1 + 2gB(q)}, \quad (30)$$

where the two accounts for the exchange terms. The denominator at small  $q$ , from Eq. (26), is given by

$$1 + 2gB(q) = 1 + \frac{32a}{\lambda^2 k_c} [\ln(k_c/q) + c]. \quad (31)$$

Since  $k_c \sim a/\lambda^2$ , the correction is of order unity, and serves to modify the spectrum, recalculated from Eq. (25) with (31), from  $k^{3/2}$  to  $k^{2-\eta}$ , with [14]  $\eta \approx (1/2) - 1/(2c + k_c\lambda^2/16a) \approx 0.36$ .

To estimate  $J^*$ , we calculate  $\Sigma(0)$  from Eq. (11) with the  $3/2$  spectrum and the leading log in  $B(q)$ , Eq. (26), and neglect the contribution for  $q > k_c$ . Then  $\Sigma(0) \approx -\kappa^2 a^2/2m\lambda^4$ , and  $\sigma(0, J) \approx -\kappa^2 J^2$ , so that at  $T_c$ ,  $J^* \approx 1/\kappa = 3/(32\sqrt{2} + 3c)$ . The self-consistent solution of Eq. (23) yields [14]  $J^* \approx 0.07$ .

The modification at  $T_c$  of the spectrum of particles at low momenta should have direct experimental consequences in

trapped Bose condensates. While a  $k^2/2m$  particle spectrum yields a flat distribution  $v^2 dn/dv$  of velocities, a more rapidly rising spectrum, e.g., the  $k^{3/2}$  discussed here, depletes the number of low momentum particles. These effects become more pronounced with a larger number of particles and flatter traps, as the level spacing ceases to control the low-energy behavior.

This work was facilitated by the cooperative agreement between the University of Illinois at Urbana-Champaign and the Centre National de la Recherche Scientifique, and also supported in part by National Science Foundation Grants No. PHY94-21309 and No. PHY98-00978, and the Deutscher Akademischer Austauschdienst. L. K. B. and L. P. S. are Unités Associées au CNRS UMR 8552 and UMR 8551.

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